

FERMI DIRAC GASSES

Fermions: Are particles of half-integer spin that obey Fermi-Dirac statistics. Fermions obey the Pauli exclusion principle, which prohibits the occupancy of an available quantum state by more than one particle.

Ideal fermion gas: Consisting of N non-interacting and indistinguishable fermions in a container of volume V held at absolute temperature T .

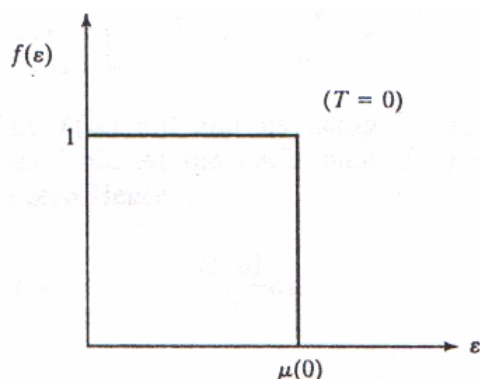
Fermi-Dirac distribution: For an ideal FD gas (non-interactions between the indistinguishable particles) of N molecules in a volume V , the most probable number of particles with ϵ_i energy is:

$$f(\epsilon_i) = \frac{n_i^*}{g_i} = \frac{1}{e^{\beta(\epsilon_i - \mu)} + 1}$$

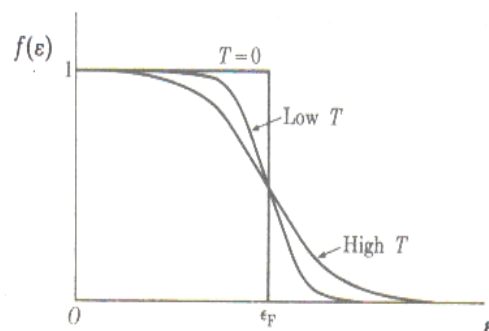
and the continuum is:

$$f(\epsilon) = \frac{N(\epsilon)}{g(\epsilon)} = \frac{1}{e^y + 1}, \quad y = \beta(\epsilon - \mu)$$

Fermi function gives the probability that a single particle state ϵ_i will be occupied by a fermion. Clearly, $0 \leq f(\epsilon) \leq 1$.



The Fermi function at $T = 0$.



Distribution function $f(\epsilon)$ at three temperatures in Fermi-Dirac statistics.

Notes:

- 1- μ no need to be negative, due to the +1 in the denominator. μ may be positive or negative.
- 2- If $\mu\beta \ll 0$, then $e^{\beta(\epsilon - \mu)} \gg 1$, and $f(\epsilon)$ reduces to the Maxwell-Boltzmann distribution.
- 3- If $\mu\beta \gg 1$
 - i- if $\epsilon \ll \mu \Rightarrow \beta(\epsilon - \mu) \ll 0 \Rightarrow f(\epsilon) = 1$
 - ii- if $\epsilon \gg \mu \Rightarrow \beta(\epsilon - \mu) \gg 0 \Rightarrow f(\epsilon) = e^{-\beta(\epsilon - \mu)}$ and fall off exponentially like Maxwell-Boltzmann distribution.

iii- if $\varepsilon = \mu \Rightarrow f(\varepsilon) = \frac{1}{2}$

iv- In the limit of $T \rightarrow 0$ we have sharp drop and

$$f(\varepsilon) = \begin{cases} 1 & \varepsilon < \mu_o \\ 0 & \varepsilon > \mu_o \end{cases} \Rightarrow N = \begin{cases} \int_0^{\mu_o} g(\varepsilon) d\varepsilon & T = 0 \\ \int_0^{\mu_o} g(\varepsilon) f(\varepsilon) d\varepsilon & T \neq 0 \end{cases}$$

Where

$$g(\varepsilon) = g_s \left(\frac{2\pi V}{h^3} \right) (2m)^{3/2} \sqrt{\varepsilon}, \quad g_s = 2s + 1$$

$$= 4\pi V \left(\frac{2m}{h^2} \right)^{3/2} \sqrt{\varepsilon}, \quad \text{for electrons}$$

Exclusion principle implies that a FD gas has a large mean energy even at absolute zero, $0 < \varepsilon < \mu_o (\equiv \varepsilon_f(0))$

$T = \begin{cases} 0 \\ T \ll T_f \\ T \approx T_f \\ T \gg T_f \end{cases}$	$\mu_o > 0, \varepsilon < \mu_o$	Very low temperature	Completely degenerate
	$\mu_o > 0$	Low temperature	degenerate
	$\mu_o = 0$	Intermediate temperature	Slightly degenerate
	$\mu < 0$	High temperature	Classical limit

At absolute zero, due to exclusion principle, all the states with $0 < \varepsilon \leq \mu_o (\equiv \varepsilon_f(0))$ are completely filled and all the states with $\varepsilon > \mu_o$ are completely empty.

Completely Degenerate Gas

Total number of particles:

$$N = \int_0^{\mu_o} g(\varepsilon) f(\varepsilon) d\varepsilon$$

$$= 4\pi V \left(\frac{2m}{h^2} \right)^{3/2} \int_0^{\mu_o} \sqrt{\varepsilon} d\varepsilon = \frac{8\pi V}{3} \left(\frac{2m}{h^2} \right)^{3/2} \mu_o^{3/2}$$

$$\Rightarrow \mu_o = \frac{h^2}{2m} \left(\frac{3N}{8\pi V} \right)^{2/3}$$

For convenience, we introduce a Fermi temperature T_f such that $\mu_o = \varepsilon_F = k_B T_f$. This can be written as:

$$T_f = \frac{\mu_o}{k_B} = \frac{h^2}{2mk_B} \left(\frac{3N}{8\pi V} \right)^{2/3}$$

Example: Metallic potassium has $\rho = 0.86 \times 10^3 \text{ kg/m}^3$ and atomic weight of $M = 39 \text{ kg/kmole}$. Find μ_o , T_f , and v_f .

Solution: We will consider one free electron per atom for monovalent atoms. Thus the concentration is:

$$\frac{N}{V} = \frac{N_a \rho}{M} = \frac{(6.02 \times 10^{26} \text{ atoms/kmole})(0.86 \times 10^3 \text{ kg/m}^3)}{39 \text{ kg/kmole}} = 1.33 \times 10^{28} \text{ atoms/m}^3,$$

$$\begin{aligned} \mu_o &= \frac{h^2}{8m} \left(\frac{3N}{\pi V} \right)^{2/3} = \frac{(hc)^2}{8mc^2} \left(\frac{3N}{\pi V} \right)^{2/3} = \frac{(12.4 \times 10^{-7} \text{ eV.m})^2}{8(0.511 \times 10^6 \text{ eV})} \left(\frac{3}{\pi} \times 1.33 \times 10^{28} \frac{\text{atoms}}{\text{m}^3} \right)^{2/3} \\ &= 2.05 \text{ eV} \end{aligned}$$

Then

$$T_f = \frac{\mu_o}{k_B} = \frac{2.05 \text{ eV}}{8.617 \times 10^{-5} \frac{\text{eV}}{\text{K}}} = 23790 \text{ K}$$

So, even at room temperature we have to treat the metallic potassium quantum mechanically.

Use

$$\begin{aligned} \mu_o &= \frac{p_f^2}{2m} \Rightarrow p_f^2 = 2m \mu_o \\ \Rightarrow v_f^2 &= \frac{2\mu_o}{m} = \frac{2\mu_o c^2}{mc^2} = \frac{2(2.05 \text{ eV}) \times (3.0 \times 10^8 \text{ m/s})^2}{(0.511 \times 10^6 \text{ eV})} = 7.22 \times 10^{11} \text{ m}^2/\text{s}^2 \\ \Rightarrow v_f &= 8.5 \times 10^5 \text{ m/s} \end{aligned}$$

Internal energy:

$$\begin{aligned} U_o &= \int_0^{\mu_o} \varepsilon g(\varepsilon) f(\varepsilon) d\varepsilon \\ &= 4\pi V \left(\frac{2m}{h^2} \right)^{3/2} \int_0^{\mu_o} \varepsilon^{3/2} d\varepsilon = \frac{8\pi V}{3} \left(\frac{2m}{h^2} \right)^{3/2} \mu_o^{5/2} \\ &= \frac{3}{5} N \mu_o \end{aligned}$$

Other thermodynamic functions are:

$$S_o = 0,$$

$$\Omega_o = -PV = U_o - S_o - \mu_o N = -\frac{2}{5} \mu_o N,$$

$$P = -\frac{\Omega_o}{V} = \frac{2}{5} \mu_o \left(\frac{N}{V} \right) = 2.71 \times 10^7 \rho \text{ atm for electrons)}$$

Thus at $T = 0$ K a fermion gas exerts a pressure. If the electrons in a metal were neutral they would exert a pressure of about 10^6 atm. The Coulomb attraction to the ions counterbalances the pressure. For $T = 0$ K the value of is positive and large.

H.W.

Prove that

$$\mu(T) \approx \mu_o \left[1 - \frac{\pi^2}{12} \left(\frac{T}{T_f} \right)^2 \right], \quad T \ll T_f$$

Notes:

- 1- $\mu(T)$ is positive for temperature below the Fermi temperature and negative for higher temperature.
- 2- As the temperature increases above, more and more of the fermions are in the excited states and the mean occupancy of the ground state falls below 1/2. In this region,

$$f(0) = \frac{1}{e^{-\beta\mu} + 1} < \frac{1}{2}$$

which implies that

$$\frac{\mu}{k_B T} < 0$$

or $\mu < 0$.

For boson gas $\mu(T)$ is negative at all temperatures and is zero at absolute zero.

AT high temperature the fermion gas approximates the classical ideal gas. In the classical limit:

$$\mu = -k_B T \ln(z) = -k_B T \ln \left[\left(\frac{2\pi m k T}{h^2} \right)^{3/2} \frac{V}{N} \right]$$

Example: for kilo-mole of the fermion ^3He gas atoms at STP, $T_f = 0.069$ K, so that $\frac{T}{T_f} = 3900$, then classically

$$\frac{\mu}{kT} = -\ln \left(\frac{z}{N} \right) = -\ln \left[\left(\frac{2\pi m k T}{h^2} \right)^{3/2} \frac{V}{N} \right] = -12.7$$

And $e^{-\beta\mu} = 3.3 \times 10^5$. The average occupancy of single particle states is very small, as in the case of an ideal gas obeying the Maxwell-Boltzmann distribution.

The thermionic emission has been worked out (see Phatria chapter 8)

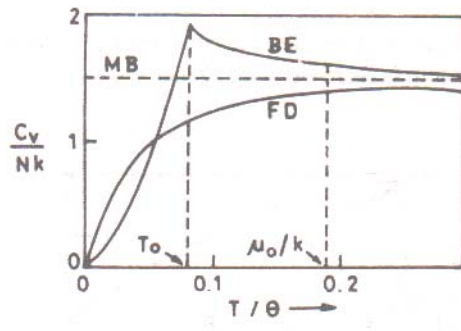


Fig. 7.5 Comparison of heat capacity of a gas according to the three statistics, $\theta = (h^2/mk)(\bar{N}/gV)^{2/3}$.