## **MULTIPLE CHOICE**

1. Ra	1. Radio telescope dishes are made as large as possible so they are (have):					
a	. better resolution	b. easier to steer c. ecc	onomical d. ea	ser to access		
2. A 4	4-m Cassegrain telescope athering or collecting light	has a hole of 55 cm in diameter ht.	. What percentage of th	e total area will remain for		
a	. 98 %	b. 99 %	c. 2 %	d. 100 %		
3. Wł	nat is the main limitation	to the resolving power of a parti	cular telescope on the su	urface of the Earth?		
a	. Light pollution	b. Seeing	c. Ozone layer	d. Color of the sky		
4. A 1	refracting telescope, com	pared with a reflecting telescope				
a c	. is free of chromatic abe . can not be used for pho	tography b. requires all d. has an obje	uminizing ective supported only fro	om its rim		
5. Wł	ny might some stars appea	ar double in blue light, though th	ney could not be resolve	d in red light?		
a b c d	<ul> <li>The wavelength of blue</li> <li>The energy of blue ligh</li> <li>The wavelength of blue</li> <li>The blue light is usually</li> </ul>	e light is longer than that of the r at is less than that of the red e light is shorter than that of the y reflected, but the red light is up	ed red sually refracted			
6. Wł	nich one of the following	criteria is not important in choo	sing a site for an optical	observatory?		
a	. Near Major Cities	b. Good Seeing	c. Dry Climate	d. Many Clear Nights		
7. The planets, which do not have moons, are:						
7. The	e planets, which do not ha	ive moons, are:				
7. The	e planets, which do not ha . Mercury and Venus	ive moons, are: b. Venus and Pluto c. Ve	enus, Pluto, and Mercury	d. Mercury and Pluto		
7. The a 8. Star	e planets, which do not ha . Mercury and Venus ndard time is based on the	ive moons, are: b. Venus and Pluto c. Ve e	enus, Pluto, and Mercury	d. Mercury and Pluto		
7. The a 8. Star	e planets, which do not ha . Mercury and Venus ndard time is based on the a. mean solar day	ave moons, are: b. Venus and Pluto c. Ve e b. sidereal month	enus, Pluto, and Mercury c. solar eclipse	y d. Mercury and Pluto d. mean lunar month		
<ol> <li>7. The</li> <li>a</li> <li>8. Star</li> <li>9. Rad</li> </ol>	e planets, which do not ha . Mercury and Venus ndard time is based on the a. mean solar day lio telescopes are always	ave moons, are: b. Venus and Pluto c. Ve e b. sidereal month reflecting telescopes because	enus, Pluto, and Mercury c. solar eclipse	y d. Mercury and Pluto d. mean lunar month		
7. The a 8. Star 9. Rac a c	e planets, which do not ha . Mercury and Venus ndard time is based on the a. mean solar day lio telescopes are always . radio waves are not ref . radio waves are not ref	ave moons, are: b. Venus and Pluto c. Ve e b. sidereal month reflecting telescopes because fracted like light waves lected like light waves	enus, Pluto, and Mercury c. solar eclipse b. radio lenses would d. you can't see radio	y d. Mercury and Pluto d. mean lunar month be impossibly large waves		
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a. vernal equinox b. autumnal equinox c. summer solstice d. winter solstice 14. The average density of Mars is

	a. 5.4 g/cm <sup>3</sup> Earth	b. 3.9 g/cm <sup>3</sup>	c. 3.4 g/cm <sup>3</sup>	d. the same as that of			
15.	15. Dark, cool regions on the photosphere are called						
	a. solar corona	b. solar wind	c. sunspots	d. terminators			
16.	The first lens the light enco	ounters in a refracting tele	scope is called the:				
	a. image	b. objective	c. prism	d. eyepiece			
17.	The Sun travels the full leng	gth of the ecliptic once ir	about				
	a. 26000 years	b. 24 hours	c. 365 days	d. 354 days			
18.	One physical characteristic	of Mercury that is appro-	ximately similar to that of the	e Earth is the:			
	a. size	b. density	c. atmosphere	d. mass			
19.	Today, astronomical image	s are usually produced ar	d recorded using				
	a. prisms	b. photosphere	c. CCDs	d. photographic plates			
20.	One of the reasons that obse	ervatories are not built or	the surface of the Moon is t	he:			
	a. transparency	b. cost	c. seeing	d. light pollution			
21.	The second lens, which is u	used to examine and mag	nify the image produced by t	he first lens, is in called the:			
	a. objective	b. primary	c. eyepiece	d. aperture			
22.	At what altitude is a star that	at is setting?					
	a. 180 °	b. 90 °	c. – 90 °	d. 0 °			
23.	What is the ratio of the light	ht gathering power of a 3	.4-m reflecting telescope to	the HST (2.4 m) telescope ?			
	a. 2.0	b. 4.0	c. 8.2	d. 1.4			
24.	What should be the smaller them in green light of 500	st separation between two nm wavelength?	o stars so the 5-m Palomar te	lescope will be able to resolve			
	a. 0.02 arc second	b. 0.01 arc second	c. 0.10 arc second	d. 0.05 arc second			
25.	The planet that has the rota	tional period approximate	ely similar to that of the Eart	h is:			
	a. the Moon	b. Venus	c. Mercury	d. Mars			
26.	The astronomical basis of the	he Lunar-Hejriah calenda	r is the:				
	a. 27 days months	b. 30 days months	c. synodic (lunar) months	d. tropical (solar) year			
27.	The astronomical basis of the	he Julian or Gregorian ca	lendars is the:				
	a. 30 days months	b. 27 days months	c. synodic (lunar) months	d. tropical (solar) year			
28.	A curved piece of glass or focus is called a:	other material that brings	all the light of a given wave	elength passing through it to a			
	a. prism	b. mirror	c. lens d.	diffraction grating			

29. Which one of the following physical properties is common between Jupiter and Saturn?

a. Rotational Period	b. Revolution Period	c. Density	d. Diameter			
30. The sidereal rotation period of the Moon is about:						
<ul><li>a. 27 times that of the Earth</li><li>c. 29 of that of the Earth</li></ul>	rth	<ul><li>b. 28 of that of the Earth</li><li>d. 30 times that of the Earth</li></ul>				
31. The number of days in the	month of Zul-Hejjah in th	ne leap years of the Lunar-Hejria	h calendar is:			
a. 27 days	b. 28 days	c. 29 days	d. 30 days			
32. Because the Earth's axis of	rotation is tilted to its ort	pit, the Earth has				
a. an atmosphere	b. a magnetic field	c. clouds	d. seasons			
33. When the side of the Moon	facing the Earth is one h	alf dark it is called a				
a. crescent moon	b. full moon	c. new moon	d. quarter moon			
34. The nearest planet to the Ea	arth is:					
a. the Moon	b. Mercury	c. Mars d. Ver	ius			
35. The mass of the Sun is appr	roximately one thousand	times that of:				
a. the Moon	b. the Earth	c. Saturn	d. Jupiter			
36. The location on the surface	of the Earth that is midw	ay between the North Pole and t	he South Pole is:			
a. on the horizon	b. on the equator	c. at Greenwich	d. near the Polaris			
37. Compared with the Earth, N	Mercury has a mass that is	S				
a. about the same	b. $\frac{1}{2}$ as much	c. $\frac{1}{4}$ as much	d. much less			
38. The number of days in the	month of February in the	simple years of the Gregorian ca	llendar is:			
a. 27 days	b. 28 days	c. 29 days	d. 30 days			
39. The tilt of the celestial equ	ator to the ecliptic is about	ut:				
a. 23½°	b. 23 <sup>h</sup> 54 <sup>m</sup>	c. 45 °	d. 67½°			
40. The winter solstice occurs	during the month of					
a. March	b. June	c. September	d. December			
41. In an equatorial mounting this axis is called the	of a telescope, one axis p	points directly at the north celest	al pole,			
a. meridian	b. zenith	c. polar axis	d. azimuth			
42. In the horizon system, the c	coordinate that is measure	ed in angular units parallel to the	horizon is the			
a. altitude	b. azimuth	c. declination	d. right ascension			
43. The rotational period of Uranus is approximately equal to that of:						
a. Mars	b. Jupiter	c. Saturn	d. Neptune			

44. C	If Dhahran, Saudi Arabia ha coordinate of about 48 ° W	as coordinates of about 5 and $16^{\circ}$ S, then Dhahran	$0^{\circ}$ E and $26^{\circ}$ N and Branch is to the	silia, Brazil has of Brasilia.			
	a. west	b. south	c. east	d. south and east			
45.	45. The summer solstice in the northern hemisphere occurs during the month of						
	a. March	b. June	c. September	d. December			
46.	The star $\beta$ Cygnus, in gener	al, should be					
	a. brighter than $\alpha$ Cygnu c. fainter than $\alpha$ Cygnus	S	<ul><li>b. a large star in Cygnus</li><li>d. a small star in Cygnus</li></ul>				
47.	The sky has been divided, b	y the International Astro	nomical Union, into 88 dist	inct areas called			
	a. asterisms	b. groups	c. constellations	d. zodiacs			
48.	What is the celestial analog	gue of latitude?					
	a. Zenith	b. Azimuth	c. Declination	d. Ecliptic			
49.	The visible disk (surface) of	f the Sun is called:					
	a. photosphere	b. solar wind	c. solar corona	d. sun spots			
50.	Longitude is measured in de	egrees of angle relative to	the				
	a. equator	b. north pole	c. vernal equinox	d. prime meridian			
51.	Magnetic storms on the Ear	th are caused by the:					
	a. solar corona	b. solar wind	c. photosphere	d. sun spots			
52.	For any location on the Eart	th there is a point directly	overhead in the sky called	the			
	a. autumnal equinox	b. vernal equinox	c. meridian	d. zenith			
53.	Relative to a sidereal day, a	solar day is					
	a. four minutes longer	b. four minutes shorter	c. the same length	d. 24 minutes longer			
54.	If you were standing at the	North Pole the stars wou	ld				
	<ul><li>a. move parallel to the hor</li><li>c. move at an oblique angle</li></ul>	izon le to the horizon	b. move perpendid d. be invisible	cular to the horizon			
55.	In general, when a telescop mounting system is called:	e is mounted in such a w	ay so it can only move alon	g the horizon and up-down, the			

a. Cassegrain	b. conventional	c. alt-azimuth	d. equatorial
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## TRUE/FALSE

56. Electronic devices are more sensitive to faint signals than photographic films.	(	)
57. The sidereal day is based on the apparent movement of the Sun while the solar day is based on the stars.	(	)
58. Refracting telescopes are more suitable in studying UV radiation than reflecting telescopes.	(	)
59. A 20-meter radio telescope studying 3-millimeter waves has poorer resolution than a 100-meter radio telescope studying 1-centimeter waves.	(	)
60. Since x-rays have such high energies and will pass right through a mirror it is not possible to use optical telescopes to study x-ray.	(	)
61. The resolution of any single radio telescope is far better than any single optical telescope.	(	)
62. The full moon is the best time to observe lunar surface features.	(	)
63. The main purpose of building observatories on high mountains is to be able to do observations in all types of the electromagnetic waves.	(	)
64. The apparent speed of the Sun changes during the year.	(	)
65. In general, at some latitudes, some stars never rise while others never set.	(	)
66. Declination is measured from celestial equator in degrees.	(	)
67. The eyepieces of refracting and reflecting telescopes can only be held from their rims.	(	)
68. It is not possible to make a large telescope from several small mirrors instead of a single large one.	(	)
69. A synodic month is shorter than a sidereal month.	(	)
70. The stars are likely to be twinkling less than the planets.	(	)
71. Today, all major observatories are built at remote sites chosen for the quality of their observing conditions.	(	)

<u>Planet</u>	<u>Radius</u> (km)	<u>Mass</u> (kg)	Ave. Distance From the Sun	Period of Revolution	Period of Rotation	Eccentricity (e)
Mercury	02439	33×10 <sup>22</sup>	057909000 km	00088 days	1408 hrs.	0.206
Venus	06052	49×10 <sup>23</sup>	0108209000 km	00225 days	5832 hrs.	0.007
Earth	06378	60×10 <sup>23</sup>	015000000 km	365.25 days	0024 hrs.	0.017
Mars	03397	64×10 <sup>22</sup>	0227939000 km	00687 days	0025 hrs.	0.093
Jupiter	71492	19×10 <sup>26</sup>	0778298000 km	04337 days	0010 hrs.	0.048
Saturn	60268	57×10 <sup>25</sup>	1429394000 km	10760 days	0011 hrs.	0.056
Uranus	25559	87×10 <sup>24</sup>	2875039000 km	30700 days	0017 hrs.	0.046
Neptune	24764	10×10 <sup>25</sup>	4504450000 km	60200 days	0016 hrs.	0.010
Pluto	0 1151	11×10 <sup>21</sup>	5915799000 km	90780 days	0153 hrs.	0.248

## Formulas and Constants (PHYS 215, term: 002)

1 hr.	=	60 min	1 min.	=	60 sec
1 day	=	24 hrs	1 week	=	7 days
1 year	=	365.25 days	1 km	=	1000 m
1 a.u.	=	$1.5 \times 10^{11} \text{ m}$	1 m	=	100 cm
1 ly	=	$9.46 \times 10^{15}$ m	1 cm	=	10 mm
1 pc	=	$3.09 \times 10^{16} \text{ m}$	с	=	$3.0 \times 10^8 \text{ m/s}$
G	=	$6.67 \times 10^{-11} \text{ m}^3/\text{kg-s}^2$	r <sub>moon-earth</sub>	=	384403 km
$M_{sun}$	=	$2.0 \times 10^{30}$ kg	R <sub>sun</sub>	=	$7.0 \times 10^5$ km
Mearth	=	$6.0 \times 10^{24}$ kg	Rearth	=	6378 km
$M_{\text{moon}}$	=	$7.4 \times 10^{22}$ kg	$\mathbf{R}_{\mathrm{moon}}$	=	1738 km
If the value of the a wavelength ( $\lambda$ ) is not mentioned in the problem then use the following: $\theta = \lambda / 50$					wing: $\theta = \lambda / 500 \times D$

 $\begin{array}{l} 1 \text{ the value of the a wavelength ($\lambda$) is not mentioned in the problem then use the following:} \\ \lambda (\text{violet}) = 400 \text{ nm} , \lambda (\text{indigo}) = 450 \text{ nm} , \lambda (\text{blue}) = 500 \text{ nm} , \lambda (\text{green}) = 550 \text{ nm} \\ \lambda (\text{yellow}) = 600 \text{ nm} , \lambda (\text{orange}) = 650 \text{ nm} , \lambda (\text{red}) = 700 \text{ nm} \end{array}$ 

 $\begin{array}{rcl} P^{2} & = & [4\pi^{2}/GM] a^{3} \\ P^{2}_{1}/P^{2}_{2} & = & a^{3}_{1}/a^{3}_{2} \end{array}$ (Kepler's **3**<sup>rd</sup> Law)  $\begin{array}{rcl} \Delta\lambda\,/\,\lambda_o &=& v\,/\,c\\ V &=& (4/3)\,\pi\,R^3\\ A &=& 4\,\pi\,R^2 \end{array}$ (Doppler effect [shift]) (volume of a sphere) (surface area of a sphere)  $= 2 \pi R$ (Circumference of a circle) С  $= \pi R^2$ а (Area of a circle) Equation of an ellipse:  $\mathbf{r} = \mathbf{a} (\mathbf{1} - \mathbf{e}^2) / (\mathbf{1} + \mathbf{e} \cos (\mathbf{\theta}))$ **2b** or  $x^2 / a^2 + y^2 / b^2 = 1$ 2a  $\mathbf{f}_1$ f<sub>2</sub> **e** is the eccentricity where:  $f_1 \& f_2$  are the locations of the two foci 2ae **a** & **b** are the semi major axis and semi minor axis, respectively

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