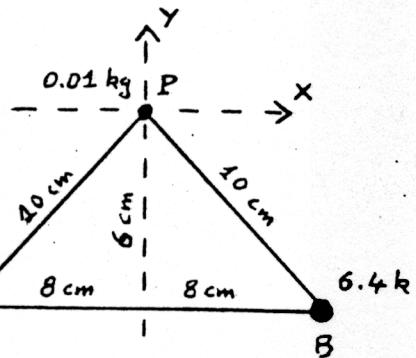
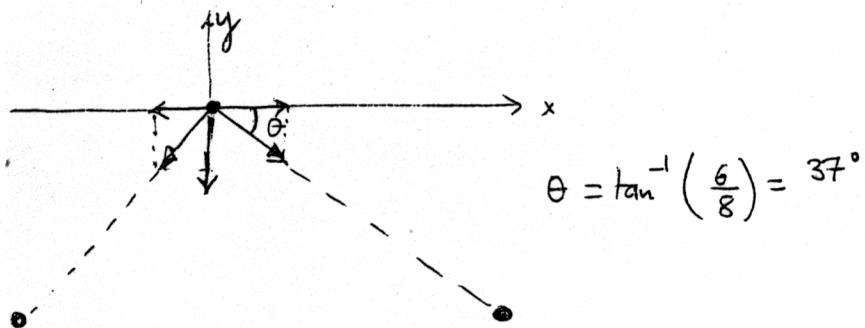


CHAPTER 14

Two spheres, each of mass 6.4 kg, are fixed at points A and B (see figure). Find the magnitude and direction of the initial acceleration of a sphere of mass 0.010 kg if released from rest at point P and acted only by forces of gravitational attraction of the spheres A and B.



- A. $0.51 \times 10^{-7} \text{ m/s}^2 (-\mathbf{j})$
- B. $0.32 \times 10^{-7} \text{ m/s}^2 (-\mathbf{j})$
- C. $0.11 \times 10^{-6} \text{ m/s}^2 (-\mathbf{j})$
- D. $0.41 \times 10^{-6} \text{ m/s}^2 (\mathbf{i} - \mathbf{j})$
- E. $0.23 \times 10^{-5} \text{ m/s}^2 (\mathbf{i} + \mathbf{j})$



$$F_x = 0$$

$$F_y = \frac{2 G m_1 m_2}{r^2} \sin 37^\circ = 5.1 \times 10^{-10} \text{ N}$$

$$\vec{F} = 0 \hat{\mathbf{i}} - 5.1 \times 10^{-10} \text{ N} \hat{\mathbf{j}}$$

$$\boxed{\vec{a} = \frac{\vec{F}}{m} = 0 \hat{\mathbf{i}} - 0.51 \times 10^{-7} \text{ m/s}^2 \hat{\mathbf{j}}}$$

A rocket is fired vertically from the earth's surface and reaches a maximum altitude above the surface of the earth equal to four earth radii. What is the initial speed of the rocket?

- A. 3.8 km/s
- B. 10 km/s
- C. 7.6 km/s
- D. 16 km/s
- E. 12 km/s

$$K_i + U_i = K_f + U_f$$

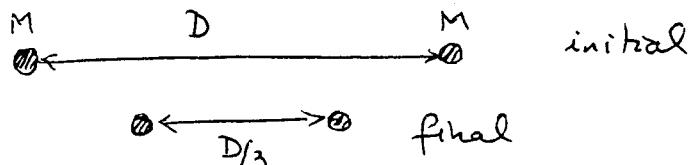
$$\frac{1}{2} m v_i^2 - \frac{GMm}{R_E} = \frac{1}{2} m v_f^2 - \frac{GMm}{5R_E}$$

$$v_i^2 = v_f^2 + \frac{8}{5} \frac{GM}{R_E}$$

$$v_i = \sqrt{\frac{8}{5} \frac{GM}{R_E}} = 10000 \text{ m/s} = \underline{\underline{10 \text{ km/s}}}$$

Two particles of mass M are initially separated by a distance D. They are released from rest and accelerate towards one another through gravitational attraction. What is the kinetic energy of each particle when their separation distance is D/3? (G = gravitational constant)

- A. $3*G*(M^{**2})/D$
- B. $G*(M^{**2})/D$
- C. $G*M/(2*(D^{**2}))$
- D. $4*G*(M^{**2})/D$
- E. $G*(M^{**2})/(2*D)$



$$K_i + U_i = K_f + U_f$$

$$-\frac{GM^2}{D} = K_f - \frac{GM^2}{D/3} = K_f - \frac{3GM^2}{D}$$

$$\Rightarrow K_f = \frac{2GM^2}{D} \quad (\text{for both particles})$$

For one particle $K_f = \frac{GM^2}{D}$

A satellite is observed to orbit a large planet close to its surface with a period of 6.00 hours. Find the average mass density of the planet. Assume that the planet is spherical.

- A. $2725 \text{ kg}/(\text{m}^{**3})$
- B. $1.29 \text{ kg}/(\text{m}^{**3})$
- C. $170 \text{ kg}/(\text{m}^{**3})$
- D. $303 \text{ kg}/(\text{m}^{**3})$
- E. $5522 \text{ kg}/(\text{m}^{**3})$

$$T^2 = \left(\frac{4\pi^2}{GM} \right) r^3$$

$$T^2 = \frac{4\pi r^3}{3} \left(\frac{3\pi}{GM} \right) = \left(\frac{V}{M} \right) \frac{3\pi}{G} = \frac{1}{P} \frac{3\pi}{G}$$

$$\Rightarrow P = \frac{3\pi}{G T^2} = 303 \text{ kg/m}^3$$

(Change to seconds)

At what altitude (in earth's radii) above the surface of the earth would the acceleration of gravity be $1/8$ of that on the surface? (R_E = radius of the earth)

- A. $0.65 * R_E$
- B. $1.83 * R_E$
- C. $2.51 * R_E$
- D. $1.02 * R_E$
- E. $0.44 * R_E$

$$a_g = \frac{GM}{(R_E + h)^2} = \frac{1}{8} \frac{GM}{R_E^2}$$

$$\Rightarrow (R_E + h)^2 = 8 R_E^2$$

$$R_E + h = \sqrt{8} R_E \Rightarrow h = 1.83 R_E$$

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$$T^2 = \left(\frac{4\pi^2}{GM} \right) r^3$$

$$T^2 = \frac{4}{3}\pi r^3 \left(\frac{3\pi}{GM} \right) = \left(\frac{V}{M} \right) \frac{3\pi}{G} = \frac{1}{P} \frac{3\pi}{G}$$

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A 100 kg spaceship is in circular orbit of radius $1.38 \times 10^{7} \text{ m}$ around the earth. How much energy is required to transfer the spaceship to a circular orbit of radius $1.92 \times 10^{7} \text{ m}$?

- A. $9.51 \times 10^{9} \text{ J}$
- B. $4.08 \times 10^{8} \text{ J}$
- C. $3.42 \times 10^{8} \text{ J}$
- D. $6.59 \times 10^{9} \text{ J}$
- E. $6.72 \times 10^{8} \text{ J}$

$$\Delta E = E_f - E_i$$

$$= -\frac{GMm}{2r_f} + \frac{GMm}{2r_i}$$

$$= + \frac{GMm}{2} \left(\frac{1}{r_i} - \frac{1}{r_f} \right)$$

$$= \frac{(6.67 \times 10^{-11})(5.98 \times 10^{24})(100)}{2} \left(\frac{1}{1.38 \times 10^7} - \frac{1}{1.92 \times 10^7} \right)$$

$$\boxed{\Delta E = 4.08 \times 10^8 \text{ J}}$$

A particle is at a height of 1000 km from the surface of the earth. Calculate the escape velocity of this particle. Assume the earth to be a perfect sphere of radius 6400 km and of mass $5.98 \times 10^{24} \text{ kg}$.

- A. 10.05 kilometers/second
- B. 11.20 kilometers/second
- C. 10.75 kilometers/second
- D. 10.38 kilometers/second
- E. 9.75 kilometers/second

$$U_i + K_i = U_f + K_f$$

$$-\frac{GMm}{r_i} + \frac{1}{2}mv_{esc}^2 = 0 \Rightarrow v_{esc} = \sqrt{\frac{2GM}{r_i}}$$

$$\boxed{v_{esc} = 10.4 \times 10^3 \text{ m/s}}$$

The planet Mars has a satellite, Phobos , which travels in a circular orbit of radius $9.40 \times 10^{10} \text{ m}$, with a period of $2.754 \times 10^4 \text{ s}$. Calculate the mass of Mars from this information.

- A. $4.56 \times 10^{26} \text{ kg}$
- B. $6.48 \times 10^{23} \text{ kg}$
- C. $3.95 \times 10^{23} \text{ kg}$
- D. $5.90 \times 10^{26} \text{ kg}$
- E. Data incomplete. Mass of Phobos is not given.

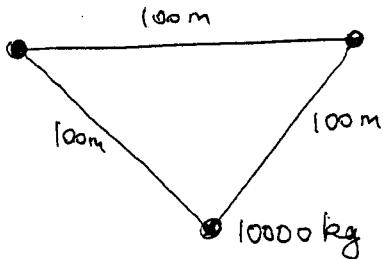
use : $T^2 = \left(\frac{4\pi^2}{GM} \right) r^3$

$$(2.754 \times 10^4)^2 = \left(\frac{4\pi^2}{6.67 \times 10^{-11} M} \right) (9.4 \times 10^6)^3$$

$$\Rightarrow M = 6.48 \times 10^{23} \text{ kg}$$

Three particles each of mass 10000 kg each are placed at the corners of an equilateral triangle with each side 100 m long. Calculate the potential energy of the system.

- A. $-2.00 \times (10^{-4}) \text{ J}$
- B. $-4.50 \times (10^{-4}) \text{ J}$
- C. $-3.18 \times (10^{-4}) \text{ J}$
- D. $-6.97 \times (10^{-4}) \text{ J}$
- E. $-8.00 \times (10^{-4}) \text{ J}$



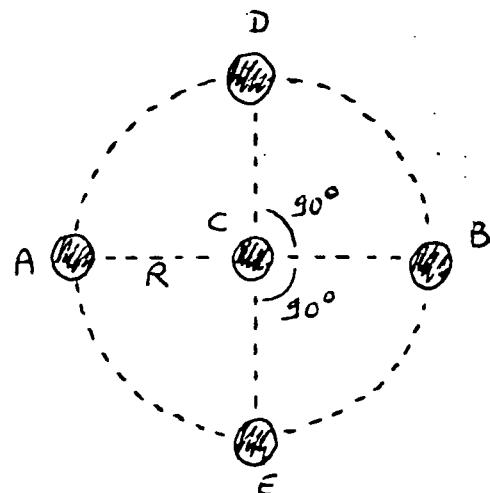
$$U = - G \frac{m_1 m_2}{r_{12}} - G \frac{m_1 m_3}{r_{13}} - G \frac{m_2 m_3}{r_{23}}$$

$$= - G \left(\frac{3m^2}{r} \right) = - 6.67 \times 10^{-11} \left(3 \times \frac{10000^2}{100} \right)$$

$$U = - 2 \times 10^{-4} \text{ J}$$

Four stars (A, B, D, E), of equal mass, rotate in the same direction around a fifth star C of the same mass located at their common center of mass (see figure). The radius of the common orbit is R. What minimum speed would star A need in order to depart from its companions for good? (express your answer in terms of G, M, R).

- A. $1.23 * (G*M/R)^{1/4}$
- B. $(G*M/R)^{1/3}$
- C. $5.32 * (G*M/R^{3/2})^{1/3}$
- D. $2.41 * (G*M/R)^{1/2}$
- E. $3.21 * (G*M/R^{3/2})^{1/2}$



$$K_i + U_i = \underbrace{K_f + U_f}_{\text{depart for good}}$$

$$\frac{1}{2} m v_i^2 - G \left(\frac{m^2}{R} + \frac{m^2}{2R} + \frac{m^2}{R\sqrt{2}} + \frac{m^2}{R\sqrt{2}} \right) = 0$$

$$\frac{1}{2} m v_i^2 = \frac{G m^2}{R} \left(1 + \frac{1}{2} + \frac{2}{\sqrt{2}} \right)$$

$$v_i = \sqrt{\frac{5.8 G m}{R}} = 2.41 \sqrt{\frac{G m}{R}}$$